

1 **Far-reaching transient motions after Mojave earthquakes require broad**
2 **mantle flow beneath a strong crust**

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12 **Abstract**

13 Geodetically observed postseismic surface displacements in the 7 years following the 1999
14 Hector Mine earthquake demonstrate a previously unrecognized broad pattern of transient
15 deformation throughout southern California and into Nevada, more than 200 km from the
16 epicenter. Unlike previous postseismic observations in which trade-offs between postseismic
17 mechanisms and the depth of flow lead to non-unique solutions, this deformation pattern can
18 only be explained by viscoelastic flow in a region of the mantle 100s of km wide and below a
19 depth of 40 km. This result enables two robust conclusions regarding the nature of lithospheric
20 strength in this region to be reached: the mantle is weaker than the lower crust, and flow occurs
21 over a wide region of mantle as opposed to within a narrow shear zone beneath the fault.

22

22 **Introduction**

23 The variation of strength with depth of continental lithosphere is much debated. Depending on
24 the composition and water content of the lower crust, warm temperatures may lead to a weak
25 viscously deforming layer sandwiched between strong upper-crustal and upper-mantle layers; i.e.
26 a “jelly sandwich” structure [e.g., Chen and Molnar, 1983]. Hotter temperatures in the mantle
27 combined with a higher water content could, however, cause the upper mantle to be weaker than
28 the lower crust; i.e. a “crème brûlée” structure [Jackson, 2002; Burov and Watts, 2006]. It also
29 continues to be a question of much debate if continental deformation at depth occurs along
30 discrete, strain-weakened shear zones or is distributed in viscously deforming lower crust and
31 lithospheric mantle. Because of the difficulty to directly determine viscoelastic strength and the
32 degree of localization of deformation in the lower crust and upper mantle, there is no consensus
33 on which region contributes most to the strength of the lithosphere and how this may vary with
34 tectonic regime, crustal age or other factors.

35 A useful approach for inferring the strength of the lithosphere is to utilize earthquakes as large
36 rock deformation experiments where coseismic stress changes induce a variety of postseismic
37 responses, including afterslip, poroelastic rebound, and viscoelastic relaxation. Each of these
38 mechanisms is capable of inducing observable postseismic surface deformation that can
39 constrain numerical models to help understand the rheological properties of the lithosphere.
40 Given the limited spatial and temporal resolution of postseismic observations, however, it has
41 proven difficult to sort out the relative contributions of each mechanism, let alone to determine
42 how viscosity varies as a function of depth. Consider the interpretation of postseismic
43 deformation following the 1992 M7.3 Landers earthquake in the Mojave Desert. Studies have
44 inferred only afterslip [Shen et al., 1994], only viscoelastic relaxation in the lower crust [Deng et

45 al., 1998]; viscoelastic relaxation predominately in the upper mantle [Pollitz et al., 2000; Freed
46 and Bürgmann, 2004], a combination of poroelastic rebound and afterslip [Peltzer et al., 1998;
47 Fialko, 2004a], or a combination of poroelastic rebound and viscoelastic relaxation in the lower
48 crust [Masterlark and Wang, 2002]. The various conclusions of these studies were influenced by
49 the use of different data sets and modeling approaches, though it is unlikely that consensus
50 would have been achieved with more comprehensive analyses, as the resolution of the post-
51 Landers deformation remains a limiting factor.

52 The Landers earthquake was followed soon after by the nearby 1999 M7.1 Hector Mine quake
53 (Figure 1). As the Hector Mine earthquake was of similar magnitude and sense of slip as
54 Landers, and perturbed the same crust and mantle, one would expect a similar postseismic
55 response. Unlike the Landers quake, however, postseismic deformation following the Hector
56 Mine quake was recorded at an extensive array of continuous GPS stations that span a very broad
57 region of southern California and into Nevada (more than 200 km from the epicenter). This first
58 of a kind far-field view (over 4 rupture lengths) of a postseismic deformation field following a
59 strike-slip earthquake, allows us to much more uniquely determine the mechanism responsible
60 for this broad deformation pattern.

61 **Observational Constraints and Modeling Approach**

62 We rely on daily time-series from continuously operating GPS sites, most of which are part of
63 the SCIGN network that became operational between 1996 and 2001. The time-series are used to
64 estimate horizontal and vertical components of linear interseismic displacement rates, coseismic
65 offsets, and a logarithmically decaying function that represents the postseismic signal (see
66 Supporting Online Material). Seven years of cumulative transient deformation resolves a broad
67 postseismic response (Figure 1, black arrows) to the 1999 M7.1 Hector Mine earthquake, as well

68 as continued deformation from the nearby 1992 M7.3 Landers earthquake. The postseismic
69 response reaches the vicinity of Yucca Mountain, Nevada, more than 200 km from the Hector
70 Mine epicenter (Figure 1, inset).

71 To understanding the mechanism responsible for this deformation pattern we use a 3-D
72 viscoelastic finite element model of the Mojave region that incorporates both rupture surfaces
73 (Supplementary Figure S2a). We use the inferred coseismic slip distribution of Fialko [2004b]
74 for the Landers earthquake, and that of Simons et al. [2002] for the Hector Mine earthquake, as
75 well as the same layered elastic structure (Supplementary Figures S2b and S2c). For every
76 candidate rheology investigated (except poroelastic rebound for which we calculate an
77 immediate cumulative response), we first simulate the Landers rupture, allow the rheology to
78 respond to these stress changes for 7 years, then simulate the Hector Mine rupture and allow the
79 rheology to respond for another 7 years. Calculated cumulative postseismic model displacements
80 over this latter 7-year period are compared to those observed to test each rheology.

81 **Results**

82 We first consider models with a layered viscoelastic structure and seek to understand the
83 depth of flow required to explain both far-field (i.e., in the region of Yucca Mountain; inset of
84 Figure 1) and mid-field (remaining stations in Figure 1) surface displacements. Since shallow
85 mechanisms, such as afterslip and poroelastic rebound [Jacobs et al., 2002, Fialko, 2004a]
86 contribute to near-field deformation (within 30 km), we do not consider near-field displacements
87 (Supplementary Figure S3) in best-fit calculations. In an initial sensitivity study, we allow
88 viscoelastic flow to occur in only one narrow depth interval at a time, and solve for the viscosity
89 required at each interval to best fit the observations based on a weighted sum of squared
90 residuals (WSSR; Figure 2a). We find that viscoelastic relaxation in the lower crust (20-28 km

91 depth) or uppermost mantle (28-40 km depth) leads to significant misfit, especially for far-field
92 motions, compared to viscoelastic flow below a depth of 40 km, with misfit minimized between
93 the depths of 40-56 km.

94 The misfit induced by viscoelastic flow in the lower crust or uppermost mantle is large
95 because flow in these depths leads to a wavelength of surface deformation that is shorter than the
96 observed broad pattern. Considering a model of lower crustal flow (20-28 km depth) with a best-
97 fit viscosity of 1.2×10^{18} Pa s. Figure 1 (blue arrows) shows that this model reasonably predicts
98 mid-field displacements, but greatly underpredicts far-field displacements in the Yucca
99 Mountain region (residual displacements are shown in Supplementary Figure S4b). In addition,
100 deformation predicted by lower crustal flow greatly underpredicts displacements along the apex
101 of curvature (green/black dashed line in Figure 1).

102 We considered a wide range of possible viscoelastic structures ranging from viscosity being
103 uniform with depth below 20 km depth, to structures where viscosity decreases rapidly with
104 depth (as might be expected due to increasing temperatures). Figure 2b shows a sample of tested
105 viscosity structures along with calculated misfits. The best models (cyan and blue lines) are those
106 where the viscosity below 40 km depth is an order of magnitude or more lower than the viscosity
107 of mantle above and two orders of magnitude less than the viscosity of the lower crust. A model
108 where flow occurs primarily below 63 km depth (black line) begins to introduce greater misfit, as
109 below this depth coseismic stress changes are too small to drive significant flow. Displacements
110 predicted by one of the best-fit models (blue line in Figure 2b), are shown in Figure 1 (red
111 arrows; residual displacements are shown in Supplementary Figure S4a). It is particularly
112 impressive how well this upper mantle flow model predicts the displacements in the area of
113 Yucca Mountain (even the rotation of azimuth between northern and southern stations), while

114 also matching the trend of deformation throughout southern California. It is worth noting that the
115 best-fit model does not require consideration of lateral variations in viscosity structure despite
116 the fact that the region encompasses several tectonic provinces. This may indicate that
117 heterogeneities in crustal properties are not mirrored by heterogeneities in the mantle beneath,
118 perhaps because the latter is mobile. The superiority of the upper mantle flow model compared to
119 that of flow in the lower crust is also evident with observed vertical displacements (Figure 3).

120 It should be noted that the viscosities shown in Figure 2 are average values over the 7 year
121 time period that lead to the best fit with respect to the observed cumulative displacements. The
122 Newtonian rheology used here cannot explain very rapid early postseismic displacements. Such
123 behavior requires a rheology where effective viscosity increases with time, such as a Burgers
124 [Pollitz, 2003] or power-law rheology [Freed and Bürgmann, 2004]. We experimented with a
125 power-law rheology and found that though calculated postseismic displacement time-series
126 evolve much differently from those resulting from Newtonian rheology, they produce similar
127 surface deformation patterns (when the depth of flow is similar), and these patterns do not vary
128 much with time.

129 We can also rule out significant contributions of localized afterslip below the seismogenic
130 zone to far-field postseismic deformation. We modeled afterslip by creating 3-km-wide shear
131 zones in the mesh that extend downward from the base of the seismogenic zone through to the
132 bottom of the model beneath both rupture surfaces. These zones extend to the north and south
133 several hundreds of kilometers (green/black dashed lines in Supplementary Figure S5), a likely
134 overestimation of the lateral extent of such shear zones. The relaxation of coseismic stresses
135 within the shear zones is controlled by a viscoelastic rheology, with all volumes outside of this
136 zone modeled as elastic. The maximum afterslip that can occur in the lower crust associated with

137 the release of coseismic stresses is simulated by assigning a very low viscosity (10^{17} Pa s) to the
138 shear zone between 20 and 28 km depth, which leads to complete relaxation of this region in the
139 7 year time frame of the postseismic observations. This model leads to surface displacements
140 that are not significant beyond about 50 km from the rupture surfaces (Figure S5, blue arrows).

141 Similarly, we can simulate the complete release of coseismic stress in a mantle shear zone
142 below a depth of 28 km. This model leads to modest displacements in the far-field, ~20% of the
143 displacements observed in the Yucca Mountain region (Figure S5, red arrows). Finally, we
144 consider narrow shear zones that cut both the lower crust and mantle beneath the Landers and
145 Hector Mine ruptures. This model leads to far-field postseismic displacement of ~22% of that
146 observed and insignificant displacement in many mid-field locations (Figure 1, yellow arrows;
147 residual displacements are shown in Supplementary Figure S4d). Since the magnitudes of many
148 of the other mid-field displacements are matched by the shear zone model, adding a component
149 of mantle flow would lead to significant overshoot at these sites. Thus, afterslip within a
150 localized shear zone below the seismogenic crust cannot be a significant source of the observed
151 broad postseismic deformation pattern. To further quantify this result, we considered shear zones
152 in the mantle ranging up to 400 km width. Only when the shear zone was large enough to
153 incorporate the area beneath the Yucca Mountain region (~300 km width) did calculated far-field
154 displacements approach the observed magnitude. Large postseismic far-field displacements
155 beneath the Yucca Mountain region can only be explained by broad viscoelastic flow in the
156 mantle.

157 We can also rule out a significant contribution to mid- and far-field postseismic displacements
158 from poroelastic rebound. We use the same parameterization of poroelastic rebound employed
159 by Fialko [2004a] to explain InSAR images following the Landers earthquake, to calculate the

160 contribution of poroelastic rebound following the Hector Mine earthquake. We find horizontal
161 surface displacements greater than 4 mm due to poroelastic rebound to be confined to within 50
162 km of the Hector Mine earthquake (Supplementary Figure S3; residual displacements are shown
163 in Supplementary Figure S4c). Figure 3c shows that significant vertical displacements predicted
164 by the poroelastic model are also confined to very near-field regions surrounding the Hector
165 Mine rupture. While a poroelastic model does predict uplift to the southwest, it significantly
166 underpredicts the observed uplift that is concentrated just beyond the reach of this mechanism.

167 **Discussion and Conclusions**

168 Previous postseismic studies have generally concentrated on relatively near-field
169 displacements, usually with only a few observations beyond a rupture length. Such analyses have
170 generally been plagued by trade-offs between different postseismic mechanisms and trade-offs
171 between the depths at which these mechanisms operate. Like the present study, previous analyses
172 have inferred a relatively weak mantle beneath the Mojave Desert [Pollitz et al., 2000; Pollitz,
173 2003; Freed and Bürgmann, 2004], but those primarily near- and mid-field studies showed trade-
174 offs with lower crustal flow. In contrast, this analysis of broad, far-field postseismic
175 displacements observed throughout southern California and into Nevada following the Hector
176 Mine earthquake requires a fairly unique solution; that flow be deep (below 40 km) and
177 distributed across 100s of km. Specifically, there are no trade-offs to lower crustal flow or
178 narrow shear zone mechanisms or poroelastic rebound to explain significant postseismic
179 displacements observed in the Yucca Mountain region, more than 200 km from the Hector Mine
180 epicenter.

181 It is important to note that the present study does not rule out the contribution of shallow
182 afterslip and poroelastic rebound to influence postseismic displacements in the near-field, as

183 suggested by previous analyses [e.g., Peltzer et al., 1998; Jacobs et al., 2002; Fialko, 2004a]. In
184 fact, near-field displacements cannot be explained solely by mantle flow and require other
185 mechanisms being active (Supplementary Figure S4a). Near-field displacements do, however,
186 contain a component from viscoelastic flow in the upper mantle (red arrows in Supplementary
187 Figure S3). Thus, analyses that do not take into account a contribution from viscoelastic flow in
188 the mantle to near-field displacements [Fialko, 2004a; Perfettini and Avouac, 2007] are likely
189 misinterpreting the postseismic observations.

190 Our inference of a relatively weak mantle 40 km below the Mojave Desert is consistent with
191 seismic velocities in the region that suggest a thin (order 10 km) mantle lid overlying a relatively
192 hot, and likely convecting, asthenosphere [Melbourne and Helmberger, 2001]. A shallow, weak
193 mantle is also consistent with thermal models derived from seismic tomography of western
194 North America [Goes and van der Lee, 2002] and evidence of a shallow asthenosphere inferred
195 from Mojave Desert xenoliths [Farmer et al., 1995]. Our inferred viscosity structure of the crust
196 and upper mantle in western Nevada is comparable to that derived from isostatic rebound
197 patterns of Lake Lahontan shorelines (5×10^{18} Pa s mantle under a much stronger crust [Bills et
198 al., 2007]). It is also consistent with a strong crust and thin mantle lid overlying shallow
199 asthenosphere (in this case at 60 km depth) inferred from analysis of postseismic deformation
200 following the 2002 Denali, Alaska earthquake [Freed et al., 2006a, 2006b]. These findings
201 suggest that at least in some backarcs or recent backarcs, the rheology is best described as a
202 continuously strong, though thin, lithosphere overlying a weak asthenosphere, the so-called
203 “crème brûlée” model [Jackson, 2002; Burov and Watts, 2006]. Considering the broad region of
204 mantle sampled by these studies, it is possible that such a model may be appropriate for much of
205 western North America and southern Alaska [Hyndman et al., 2005; Dixon et al., 2004].

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212

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272

273 **Figure Captions**

274 **Figure 1.** Cumulative GPS observed postseismic horizontal surface displacements (transient
275 component) for the 7 year period following the 1999 Hector Mine earthquake compared to those
276 calculated by models of viscoelastic flow and afterslip within narrow shear zones. Stations
277 within 20 km of the Landers and Hector Mine rupture surfaces have been excluded from this
278 comparison (see Supplementary Figure S3 for near-field displacements). SAF: San Andreas

279 Fault. Inset: Enlargement of Yucca Mountain region. GPS errors are shown at the 95%
280 confidence level, as computed using a correlated noise model as described in Herring [2003]. See
281 Supplementary Table S1 for tabulated GPS data. Upper mantle viscosity structure is the blue line
282 in Figure 2b. Lower crustal viscosity structure is the red line shown in Figure 2b. Green/black
283 dashed line shows the apex of the curved deformation field to the southwest of the Landers
284 rupture. Transient time-series of labeled stations are shown in Supplementary Figure S1.

285 **Figure 2.** (a) Weighted misfit as a function of the depth interval (defined by thin gray lines) at
286 which viscoelastic flow is allowed to occur (i.e. this is a composite of results from 12 models of
287 flow at various depths). We quantify misfit (WSSR) as $\sqrt{(1/m) \sum (d_o - d_c)^2 / \sigma^2}$, where d_o
288 and d_c are the observed and calculated displacements, σ is the observational error, and m is the
289 total number of observations. The viscosity values were tuned to match the observed
290 displacements of “all stations” shown in Figure 1. Best-fit viscosity values decrease with depth
291 from 4.6×10^{18} Pa s for the lower crustal layer from 20-24 km depth to 4.0×10^{17} Pa s for the region
292 from 109-123 km depth. “Far-field only” refers to the misfit of these same models to just the far-
293 field Yucca Mountain region stations (inset of Figure 1). (b) Viscosity versus depth profile for a
294 variety of viscoelastic flow models considered. The viscosity of all models was tuned to match
295 the observed displacements of all stations in Figure 1. Misfits are shown for calculations based
296 on all stations (All) and for just the far-field Yucca Mountain stations (Far-field).

297 **Figure 3.** Cumulative GPS observed postseismic vertical surface displacements (white bars show
298 uplift, black bars show subsidence) for the 7 year period following the 1999 Hector Mine
299 earthquake compared to those calculated (contours) by models of viscoelastic flow in (a) the
300 upper mantle (blue line viscosity structure in Figure 2b) and (b) the lower crust (red line in
301 Figure 2b) and (c) poroelastic rebound. GPS errors are shown at the 68% confidence level.

302 Stations with estimated errors greater than 5 mm in the 7-year span of the observations have been
303 excluded from this comparison.

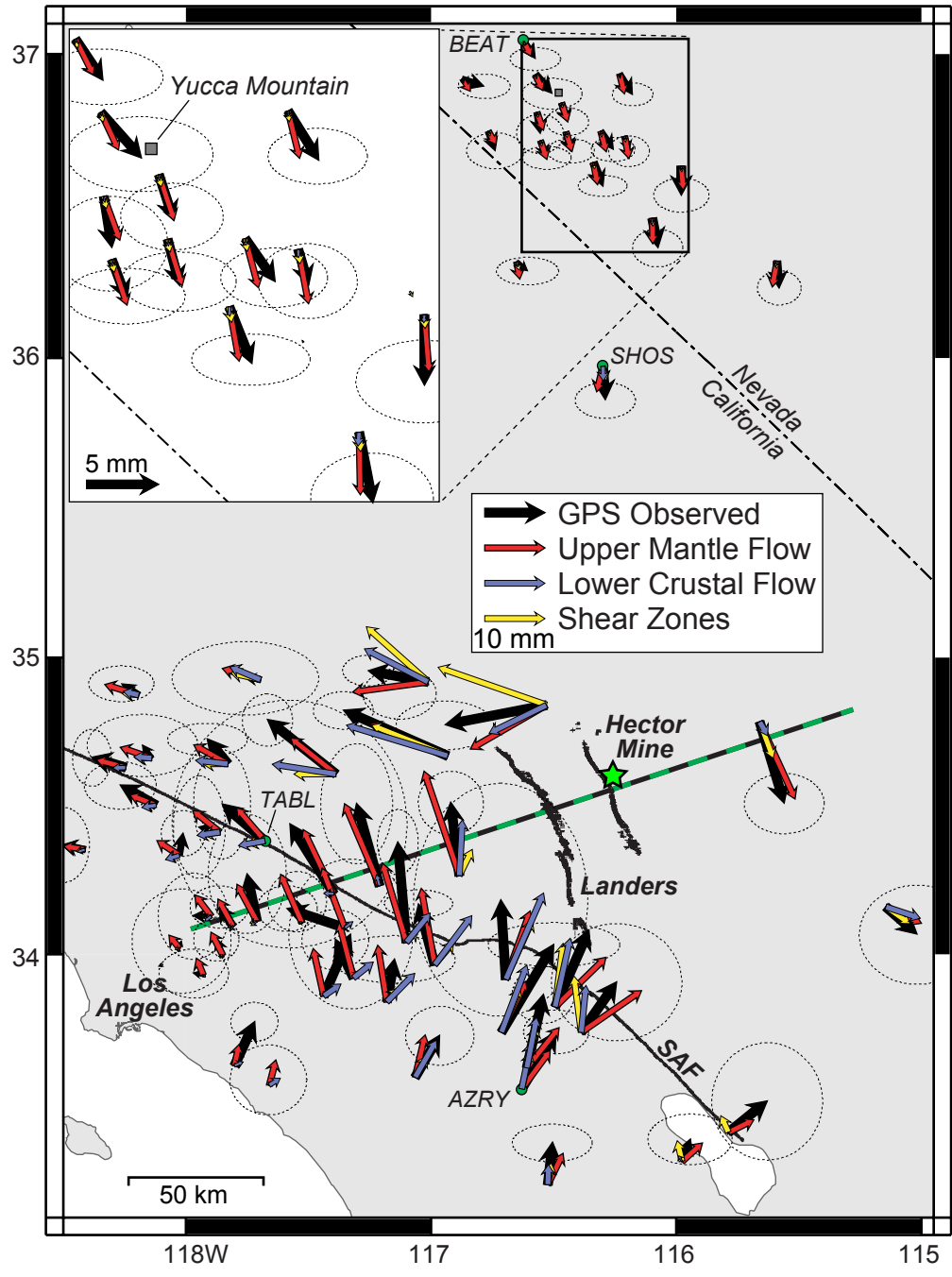


Figure 1

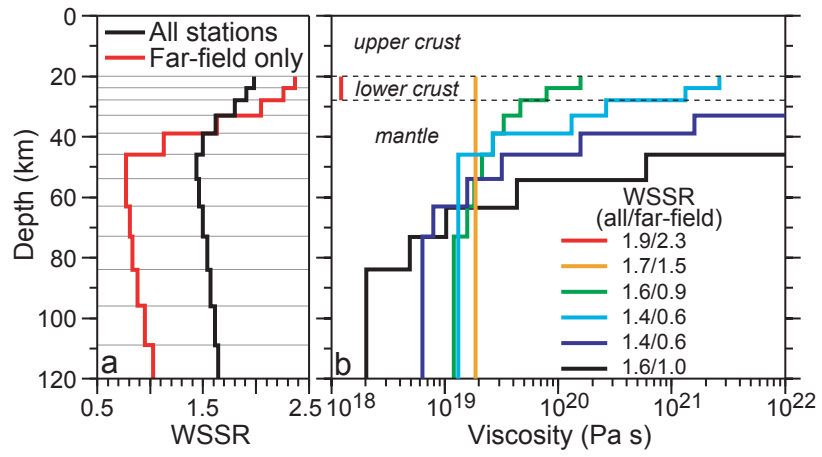


Figure 2

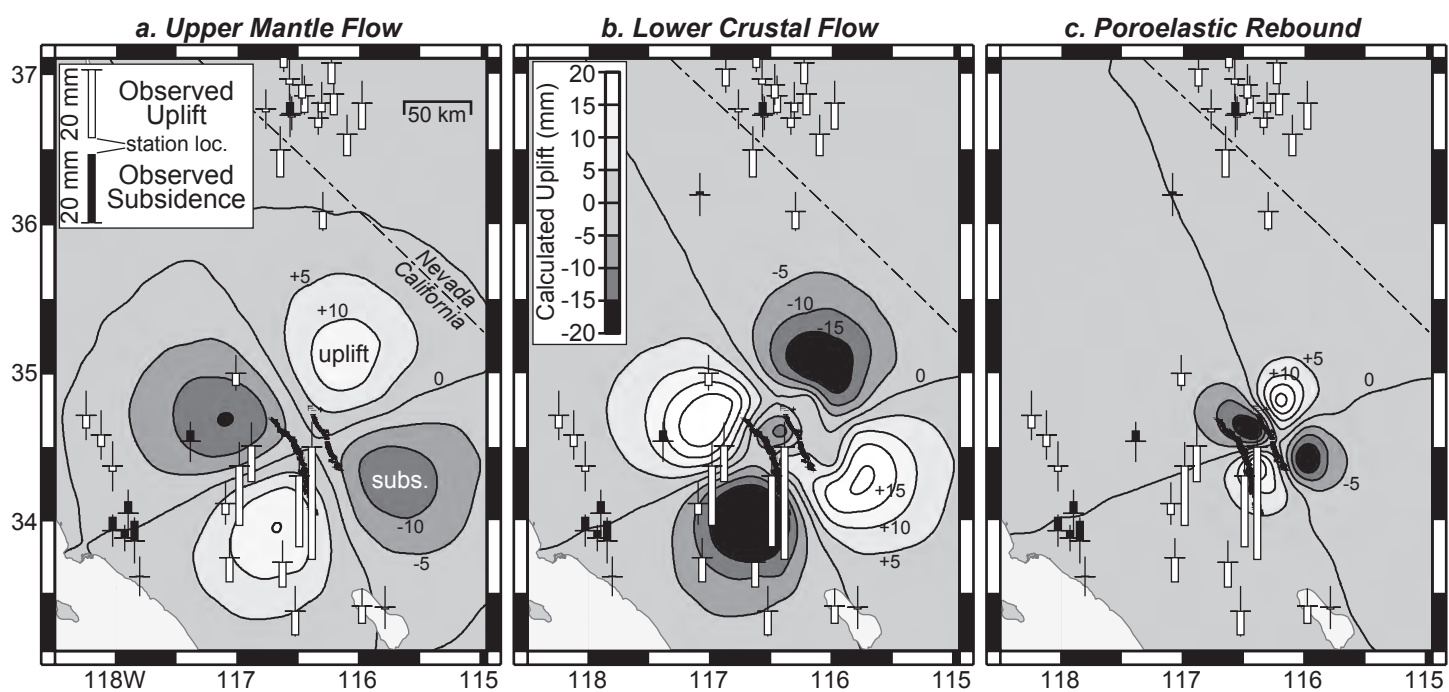


Figure 3

Supplementary Material: GPS Processing

GPS Time series were generated with the Gamit/Globk software package [Herring et al., 2006] from loosely constrained GPS position and orbit products (gamit h-files) generated at the SOPAC analysis center [<http://sopac.ucsd.edu>]. The North America Reference frame used was realized using 124 GPS sites, frame sites, distributed across North America, Greenland and Hawaii. For each day, the GPS coordinate system was aligned to the linear motion model of these frame sites through rotation, translation and scale. None of the frame sites were closer than 430 km from the epicenter of the Hector Mine earthquake.

The full time-series were used to estimate horizontal and vertical components of linear interseismic displacement rates, coseismic offsets on the day of the Hector Mine earthquake, and a decaying relaxation function starting at the day of the earthquake. The postseismic data were fit to a natural logarithmic function of the form,

$$X(t) = x_0 + v(t - t_m) + [C + \lambda \ln(1 + (t - t_{eq})/\tau)]\Theta(t_{eq}), \quad (1)$$

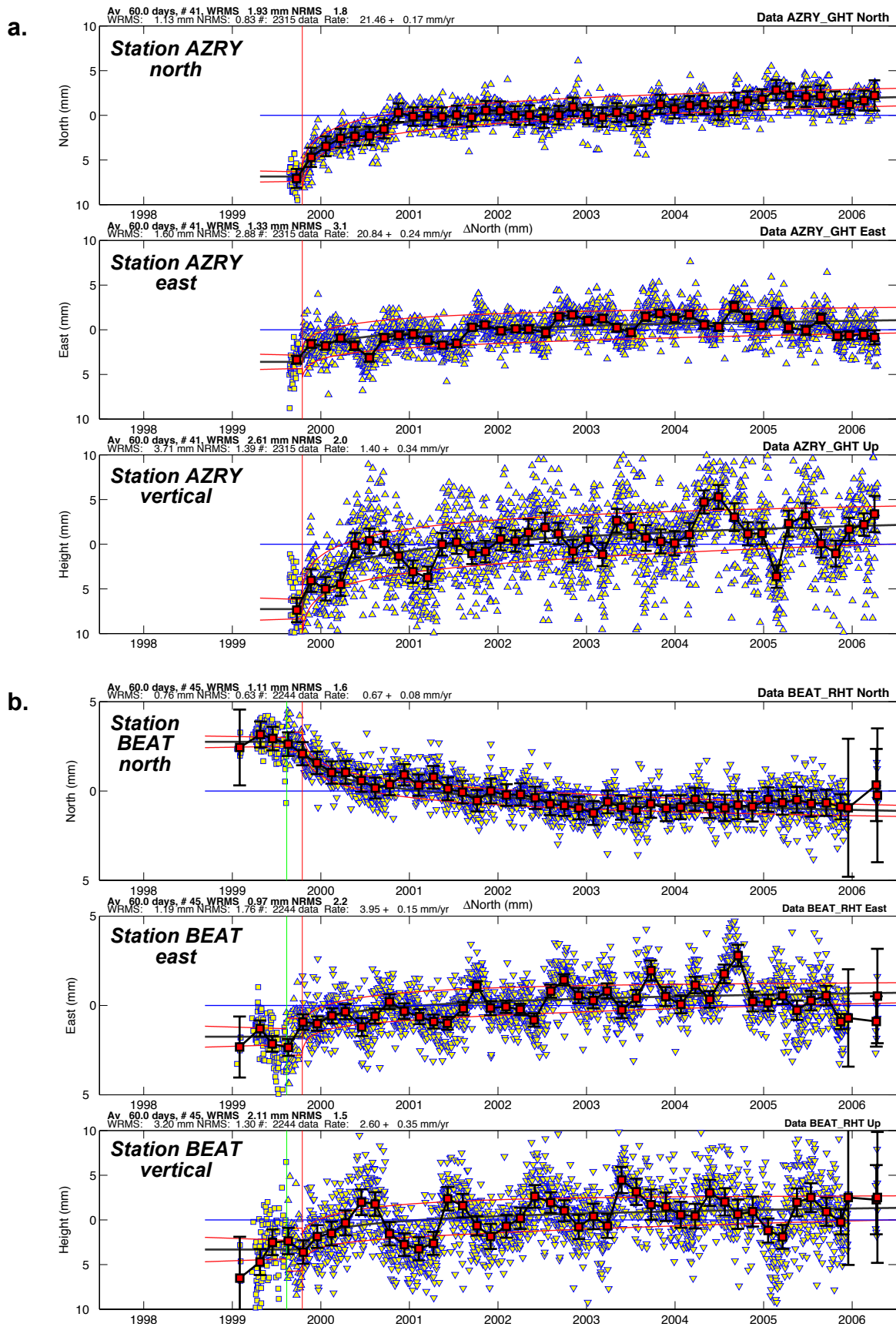
where v is the linear rate, C the coseismic offset, $\Theta(t_{eq})$ is a Heavy-side step function at the time of the earthquake, λ is the amplitude of the logarithmic term, τ is the time constant of the logarithmic form used for all the time-series (here 10 days), and t , t_m , and t_{eq} are the time of each daily epoch, of the first epoch and of the earthquake, respectively (<http://gpsweb.mit.edu/~tah/GGMatlab>). Thus, the magnitude of postseismic displacements at any epoch (e.g., 7 years as chosen for the postseismic analysis) can be estimated from the optimal log-coefficients fit to the individual time-series. Examples of transient time-series in the north, east, and vertical directions are shown in Supplementary Figure S1. Tabulated GPS data is shown in Supplementary Table S1. It should be noted that, according to our finite element model results, about 1/3 of the cumulative displacements (less for near-field stations) arises from continued viscoelastic relaxation following the 1992 Landers earthquake. Displacements from two separate earthquakes can be modeled by a single logarithmic curve and a linear rate because the contribution from the Landers earthquake from 7 to 14 years after its occurrence has a slowly varying rate. As noted in the Hector Mine time-series shown in Supplementary Figure S1, very fast rate changes only occur in the first few years after the events.

We ran tests that show that the changing the time constant does affect the estimate of the log coefficients, but when the total displacement over 7 years is computed (the quantity we are comparing in this analysis), the estimated total offsets are very close (<1 mm differences for 10days compared with 100 days). If the time constant is increased to 1000 days, the fit to the initial (first year) transients degrades, though the total offset estimate is still within 2-3 mm. The uncertainty of the estimates increases with longer time constants.

Vertical GPS displacements are inherently more noisy and susceptible to non-tectonic influences. We thus restrict our comparisons (Figure 3) to stations with estimated error of less than 5 mm over the 7-year observational period.

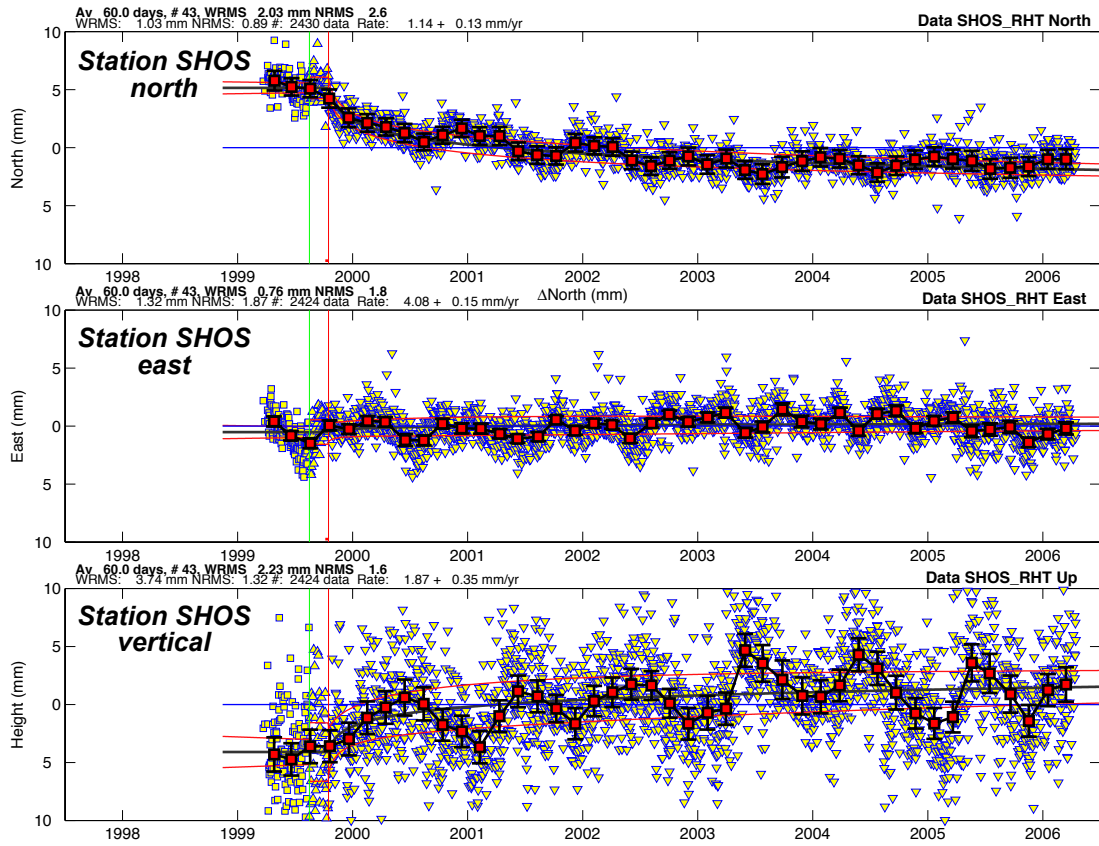
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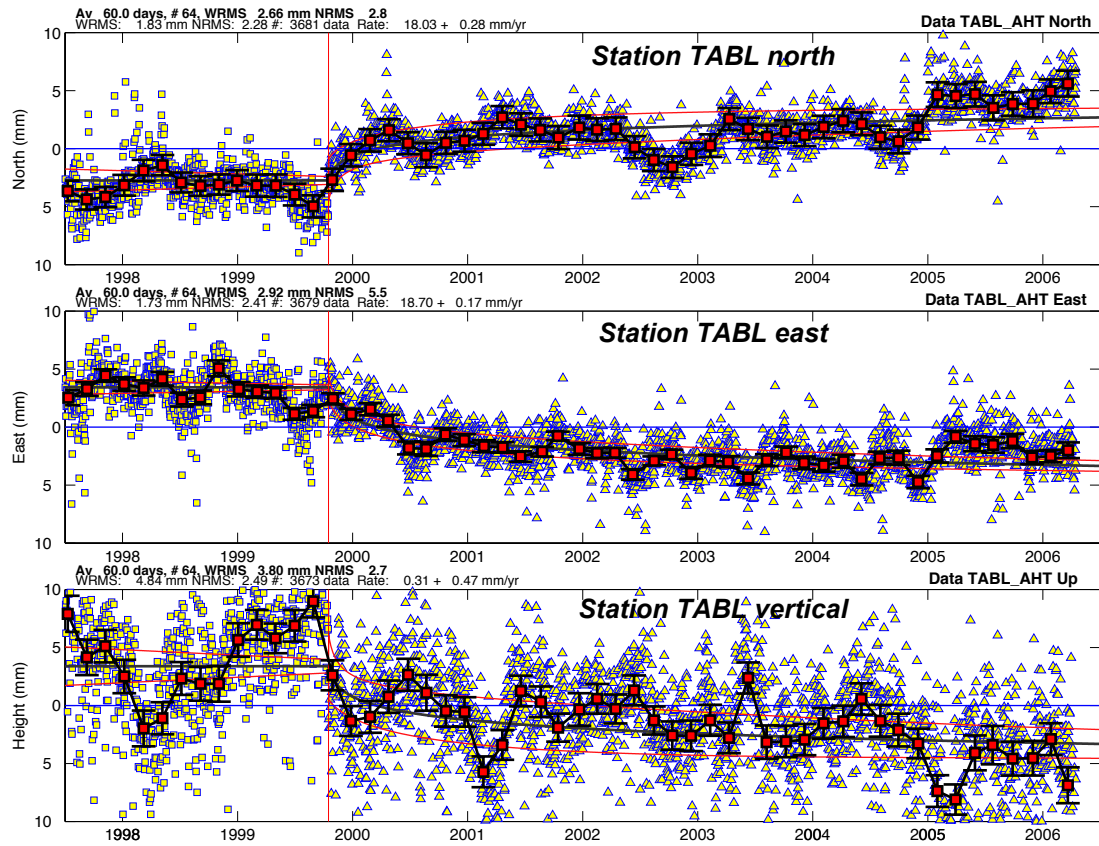


Supplementary Figure S1. Example transient GPS time-series (i.e. only the postseismic component) for four stations (a) AZRY, (b) BEAT, (c) SHOS, and (d) TABL. The secular background (linear) trend has been removed from each time-series. The location of these stations are indicated in Figure 1. Vertical red lines show the time of the Hector Mine earthquake (vertical green lines show other estimated offsets [removed from the plot] due to either an antenna or radome change). The blue horizontal line is the arbitrary zero datum. The red squares with error bars are 60-day averages with 1-sigma error bars computed from the rms scatter within the 60-days. The thick black line that tracks through the data is the logarithmic fit, with the thin red lines on either side showing ± 1 sigma errors.

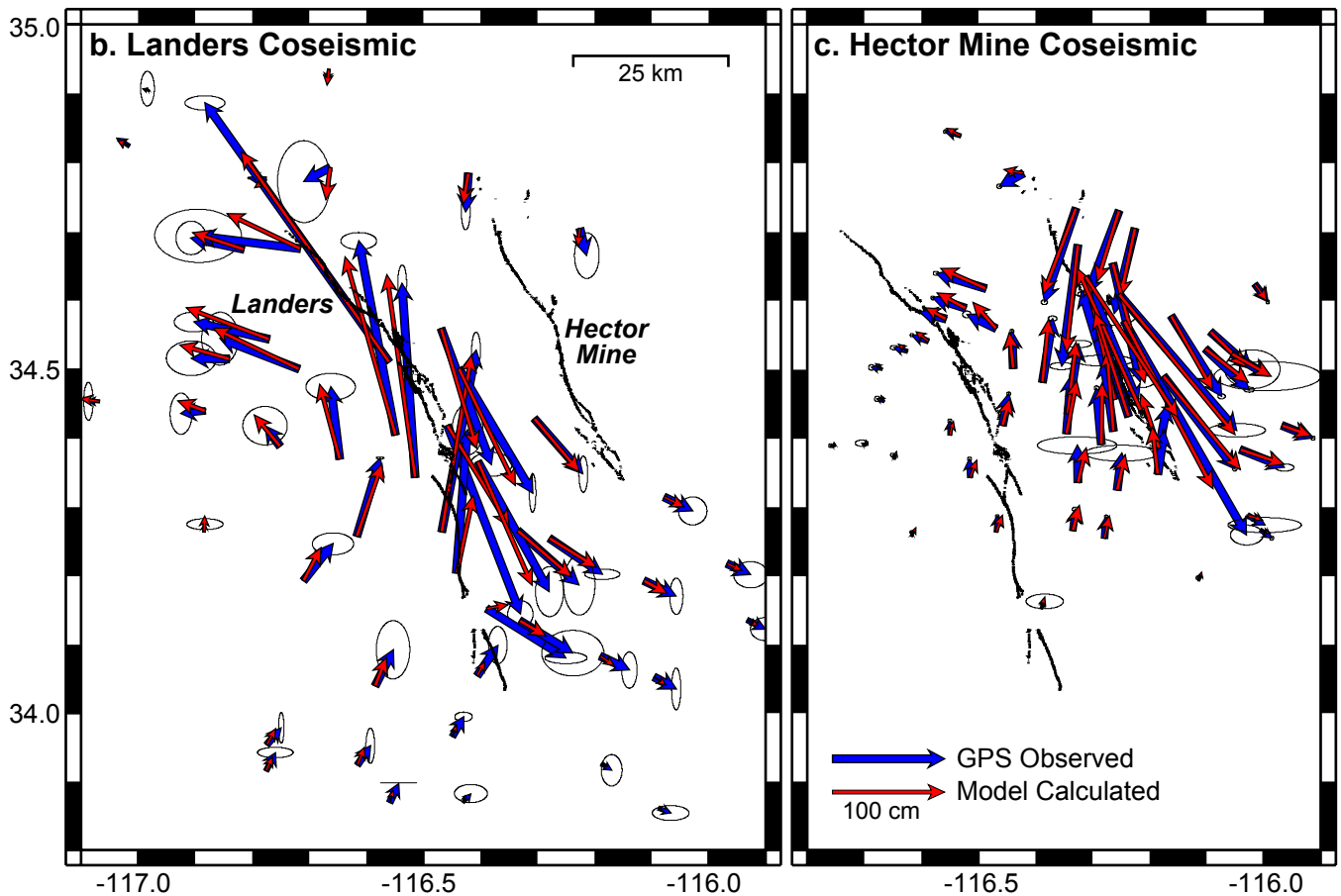
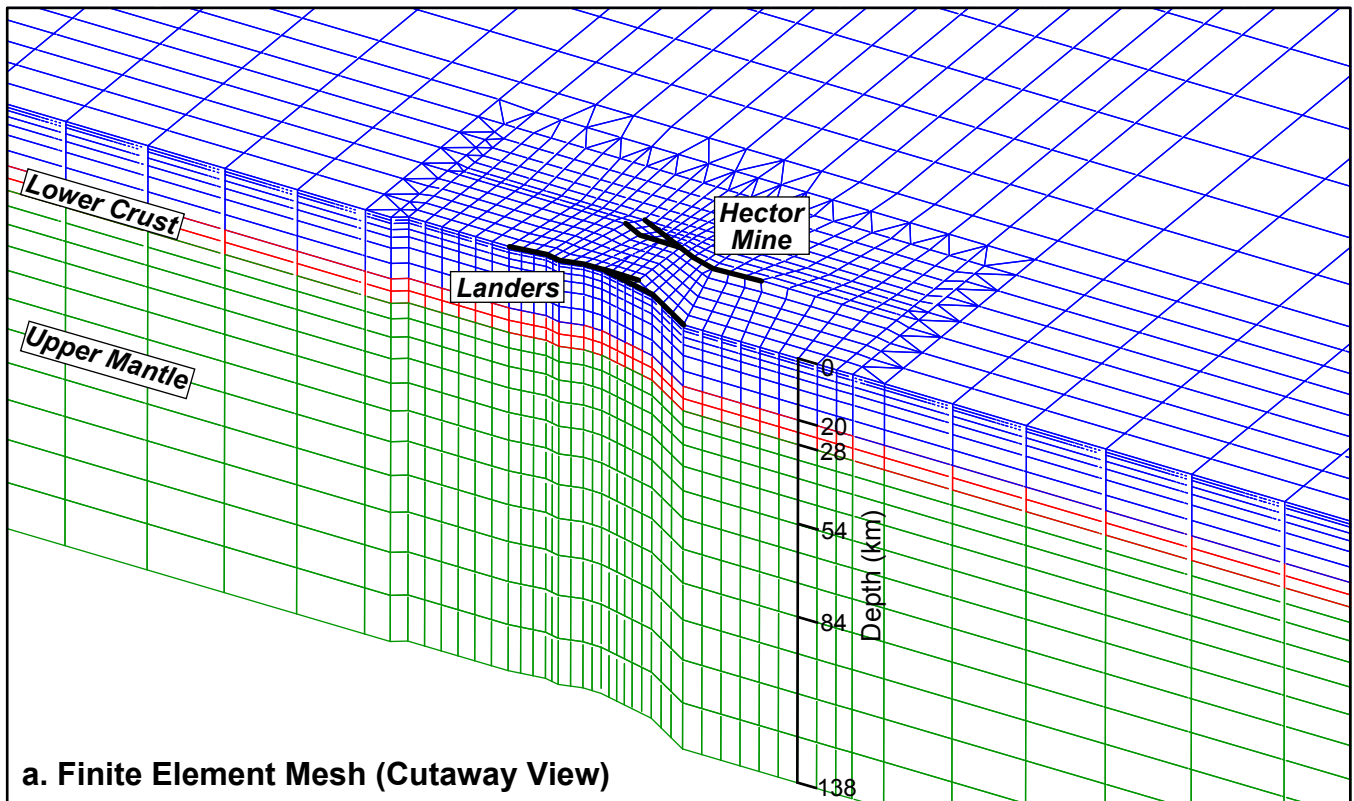
C.



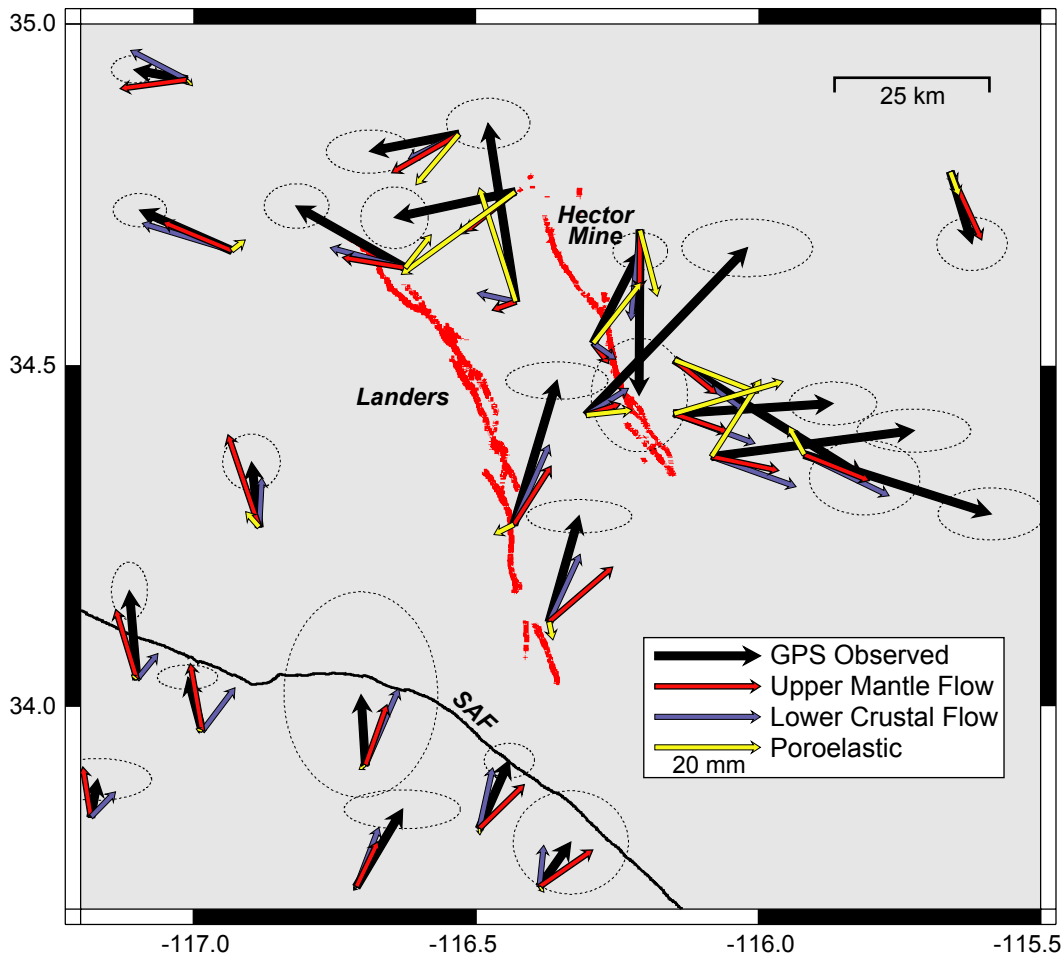
d.



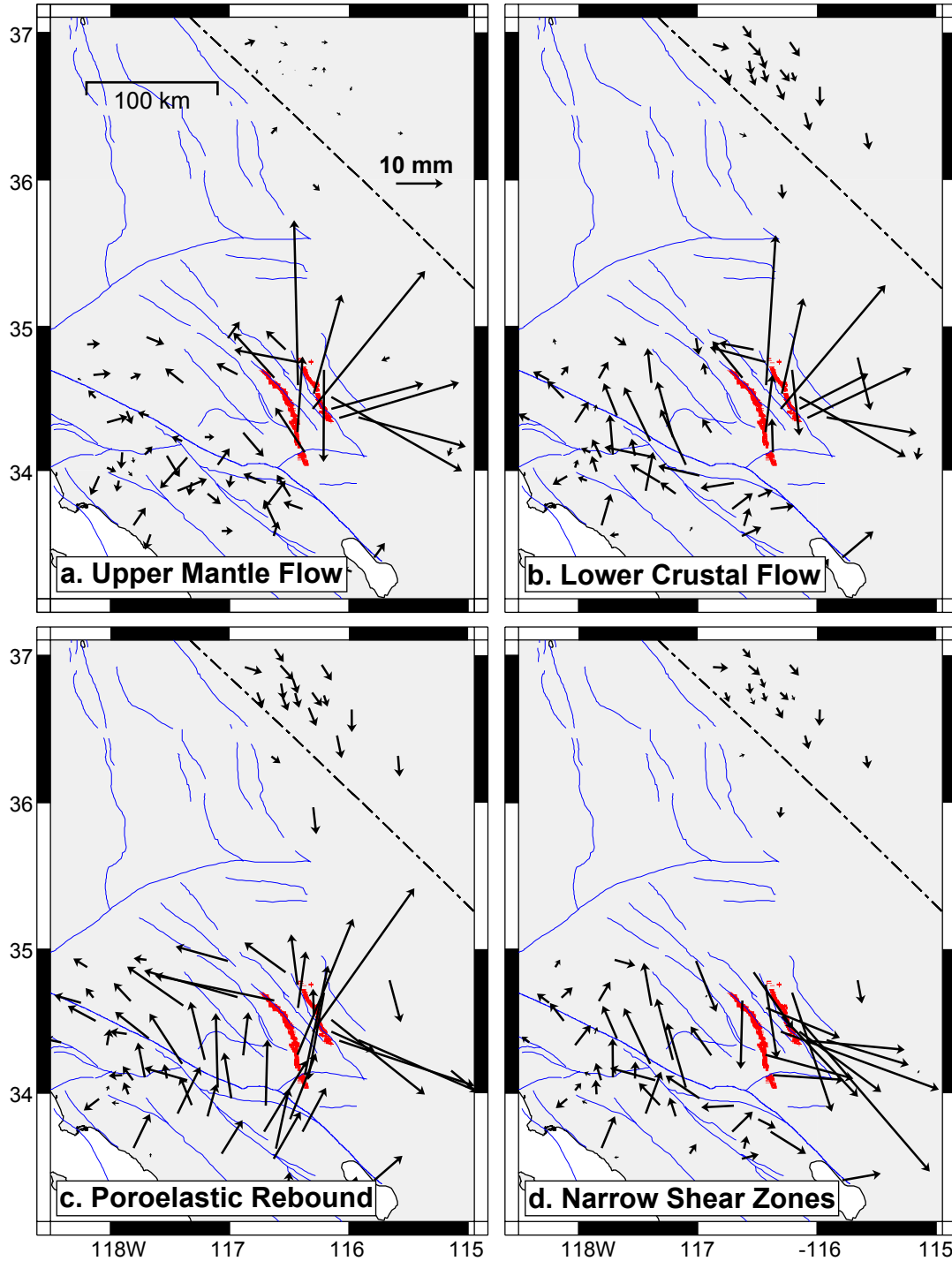
Supplementary Figure S1 (cont.)



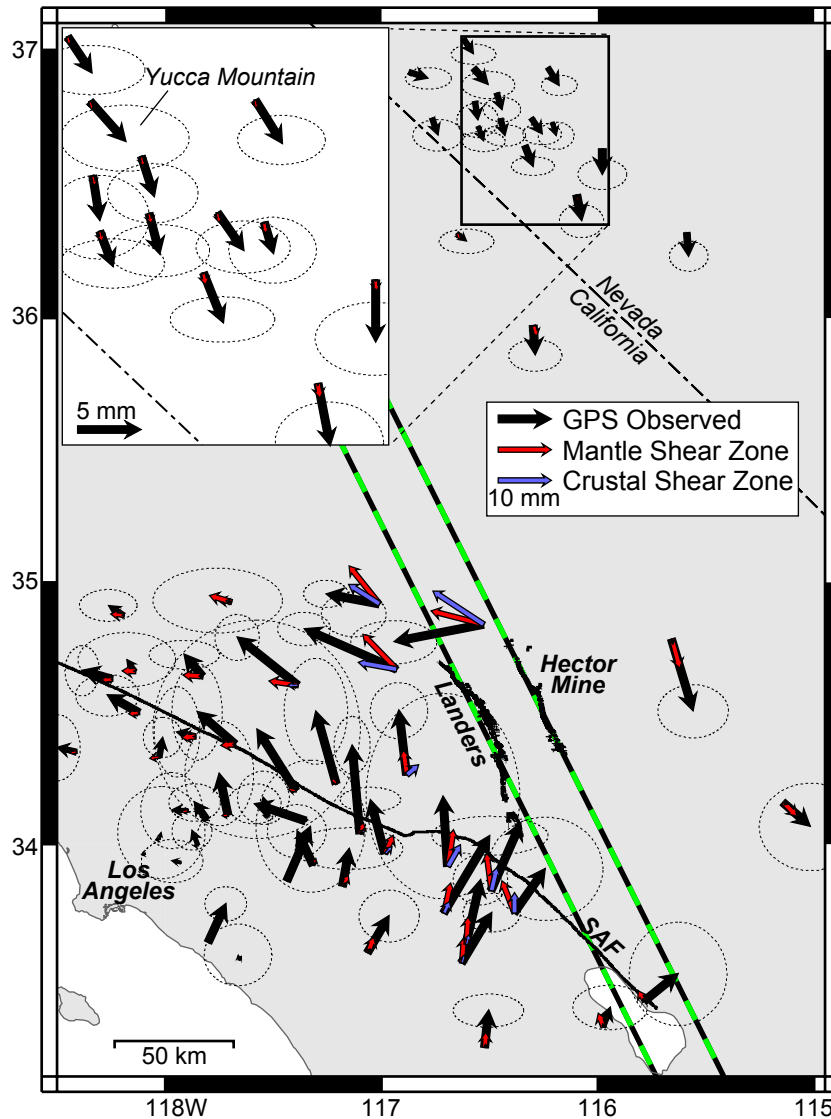
Supplementary Figure S2. (a) Cutaway view of finite element mesh along a surface through the Landers rupture surface. Surface trace of the Landers and Hector Mine rupture surfaces are shown as black lines. The mesh contains 36,000 elements and spans an area of 900 x 700 x 138 km. (b) Shows a comparison between GPS observed [Fialko, 2004b] and calculated coseismic displacements associated with the 1992 Landers earthquake. (c) Shows a comparison between GPS observed [Simons et al, 2002] and calculated coseismic displacements associated with the 1999 Hector Mine earthquake. GPS uncertainties are shown at the 95% confidence level. Black lines show the Landers and Hector Mine rupture surfaces.



Supplementary Figure S3. Cumulative near-field GPS observed postseismic horizontal surface displacements (transient component) for the 7 year period following the 1999 Hector Mine earthquake compared to those calculated by models of viscoelastic flow and poroelastic rebound. SAF: San Andreas Fault. GPS errors are shown at the 95% confidence level, as computed using a correlated noise model as described in Herring [2003]. Upper mantle viscosity structure is the blue line in Figure 3b. Lower crustal viscosity structure is the red line shown in Figure 3b. Model results generally fail to match observed GPS displacements near the faults as these are due to multiple mechanisms, including shallow afterslip, which is not included in any of these models. Note how displacements associated with poroelastic flow (yellow arrows) are confined primarily to within 10s of km of the Hector Mine rupture surface.



Supplementary Figure S4. Residual displacements (observed minus calculated) for models of (a) viscoelastic flow within the upper mantle (blue line in Figure 3b), (b) viscoelastic flow within the lower crust (red line in Figure 3b), (c) poroelastic rebound, and (d) within narrow 3-km-wide shear zones within the lower crust and upper mantle and extending 100s km to the north and south of the rupture region.



Supplementary Figure S5. Cumulative GPS observed postseismic horizontal surface displacements (transient component) for the 7 year period following the 1999 Hector Mine earthquake compared to those calculated by models of complete relaxation within narrow shear zones below the rupture surfaces and extending hundreds of kilometers to the north and south (green/black dashed lines). Stations within 20 km of the Landers and Hector Mine rupture surfaces have been excluded from this comparison. SAF: San Andreas Fault. Inset: Enlargement of Yucca Mountain region. GPS errors are shown at the 95% confidence level. Crustal shear zone is from 20 to 28 km depth. Mantle shear zone is from 28 to 120 km depth.

Supplementary Table S1. Seven year cumulative displacements and associated errors at GPS stations used in this study. * denotes near-field stations not used in model testing.

Long.	Lat.	East (mm)	North (mm)	Err_E (mm)	Err_N (mm)	Station	Long.	Lat.	East (mm)	North (mm)	Err_E (mm)	Err_N (mm)	Station
-116.429	34.594	-5.27	33.72	3.22	1.94	AGMT*	-116.099	36.459	0.94	-4.83	1.72	1.28	JOHN
-114.932	36.319	2.61	-3.49	1.83	1.50	APEX	-116.433	34.267	8.15	27.23	3.94	1.44	LDES*
-117.522	36.050	7.32	-3.77	3.99	1.83	ARGU	-116.209	34.699	-0.17	-31.17	3.72	4.38	LDSW*
-116.630	33.540	5.71	9.54	2.22	1.50	AZRY	-118.139	34.662	-1.55	2.16	3.77	2.11	LINJ
-117.897	34.126	-3.00	0.11	2.38	2.94	AZU1	-116.308	36.746	2.05	-2.94	1.50	0.83	LITT
-116.884	34.264	-1.44	12.20	2.22	2.16	BBRY	-119.104	34.734	-1.61	-0.50	2.33	1.55	LVMS
-116.621	37.040	1.89	-2.83	1.72	0.78	BEAT	-117.437	33.857	4.49	10.37	4.22	5.10	MATH
-117.065	33.578	4.10	7.04	2.33	2.00	BILL	-115.979	36.633	0.00	-4.83	1.89	1.16	MERC
-118.095	33.962	-3.33	-2.33	2.00	3.00	BKMS	-117.318	33.918	-3.00	6.43	2.27	2.05	MLFP
-114.715	33.610	4.22	2.61	2.66	5.16	BLYT	-116.422	32.892	-10.54	1.89	5.66	3.38	MONP
-116.985	33.963	-2.72	10.21	2.33	0.94	BMRY	-117.210	34.231	-3.83	13.70	2.44	4.10	MSOB
-117.012	34.919	-10.04	1.89	1.72	1.05	BSRY	-116.525	33.211	0.78	7.04	2.72	1.28	MVFD
-116.872	36.918	3.61	-1.16	1.77	0.83	BULL	-116.148	34.509	35.89	-22.24	4.27	2.88	NBPS*
-116.451	36.745	0.83	-3.16	1.61	0.83	BUST	-115.918	34.370	35.28	-11.20	4.10	2.00	OPBL*
-118.026	34.333	0.61	3.72	1.50	1.55	CHIL	-116.305	34.428	30.45	31.61	5.05	2.22	OPCL*
-116.766	36.746	1.11	-3.16	1.94	1.22	CHLO	-116.083	34.367	38.27	4.94	4.38	1.66	OPCP*
-117.828	34.641	-3.44	4.16	2.72	1.11	CHMS	-116.149	34.430	30.01	1.94	3.38	1.66	OPCX*
-116.666	31.871	4.33	2.33	2.55	1.77	CIC1	-116.292	34.533	8.87	17.42	2.11	1.44	OPRD*
-117.709	34.110	-1.66	7.93	2.66	4.05	CLAR	-117.695	34.925	-2.55	0.33	4.99	2.50	PHLB
-118.411	34.353	-3.61	0.83	1.77	2.55	CMP9	-117.243	32.665	13.42	-11.37	3.99	4.55	PLO3
-116.387	33.733	5.88	8.38	4.44	3.99	COTD	-117.182	33.836	1.28	7.15	4.16	1.61	PPBF
-116.569	36.808	0.55	-3.49	1.61	1.33	CRAT	-116.494	33.819	5.66	12.87	1.94	1.33	PSAP
-117.100	34.039	-1.44	16.75	1.39	2.22	CRFP	-117.807	34.092	-2.33	3.33	2.00	1.44	PSDM
-115.735	33.070	-0.72	0.22	5.16	4.33	CRRS	-118.245	34.629	-5.66	1.72	1.33	1.94	QHTP
-116.370	34.124	5.66	19.97	4.05	1.28	CTMS*	-116.625	34.644	-20.58	11.65	2.50	1.72	RDMT*
-115.788	33.390	6.49	5.27	3.77	3.99	DHLG	-116.554	36.715	1.05	-2.77	1.66	0.78	RELA
-116.712	33.733	8.71	14.64	4.49	1.50	DSSC	-116.468	36.840	1.00	-3.11	1.44	0.94	REPO
-117.860	34.413	-3.49	0.83	1.77	3.49	DVPB	-118.026	34.019	0.33	1.22	3.38	3.55	RHCL
-117.526	34.104	-0.83	3.61	2.00	2.05	EWPP	-116.610	33.611	2.77	12.26	2.11	1.94	ROCH
-118.894	34.410	0.00	-5.88	2.83	6.38	FMTF	-117.085	36.218	2.33	-3.49	3.27	2.61	ROGE
-117.398	34.204	-7.32	11.65	7.99	8.43	GHRP	-118.193	34.875	-2.88	1.89	2.22	1.22	RSTP
-115.660	34.784	4.10	-13.59	2.72	2.05	GMRC	-117.353	34.089	-9.60	3.22	2.05	1.55	RTHS
-116.430	34.755	-22.85	-4.88	2.61	2.38	HCMN*	-116.650	36.316	1.77	-1.44	2.11	0.94	RYAN
-115.032	32.706	3.77	-8.43	5.55	6.05	IID2	-117.661	33.553	-0.11	0.89	2.61	2.38	SBCC
-115.145	34.158	5.21	-4.55	3.94	3.38	IMPS	-117.388	34.607	-11.43	9.10	1.72	1.77	SCIA

Supplementary Table 1 (cont).

<u>Long.</u>	<u>Lat.</u>	<u>East (mm)</u>	<u>North (mm)</u>	<u>Err_E (mm)</u>	<u>Err_N (mm)</u>	<u>Station</u>
-116.696	33.913	-0.83	13.31	5.93	7.93	SGPS
-116.299	35.971	0.39	-5.60	2.05	1.22	SHOS
-117.250	32.865	2.00	13.87	5.99	9.76	SIO3
-116.211	36.730	0.72	-2.44	1.39	1.05	SKUL
-115.978	33.292	1.39	3.72	2.88	1.72	SLMS
-115.588	36.320	0.33	-4.44	1.50	1.16	SMYC
-117.929	33.927	-1.61	0.33	2.38	1.66	SNHS
-117.849	33.993	-1.11	2.83	1.55	1.50	SPMS
-115.466	31.045	3.33	1.00	7.15	4.10	SPMX
-116.338	36.645	1.55	-3.88	1.66	0.72	STRI
-117.678	34.382	-7.15	6.16	1.94	3.72	TABL
-116.574	36.932	2.77	-3.11	2.05	1.05	TATE
-116.230	36.935	2.11	-3.38	1.39	0.78	TIVA
-117.803	33.618	3.38	7.38	1.61	1.33	TRAK
-116.530	34.839	-16.92	-3.27	3.33	1.66	TROY
-116.085	33.030	8.10	-0.67	5.32	4.10	USGC
-118.121	34.502	-5.99	3.38	2.44	1.61	VNPS
-116.932	34.669	-17.42	7.65	2.05	1.28	WOMT